



Marin

Manouverability of ships

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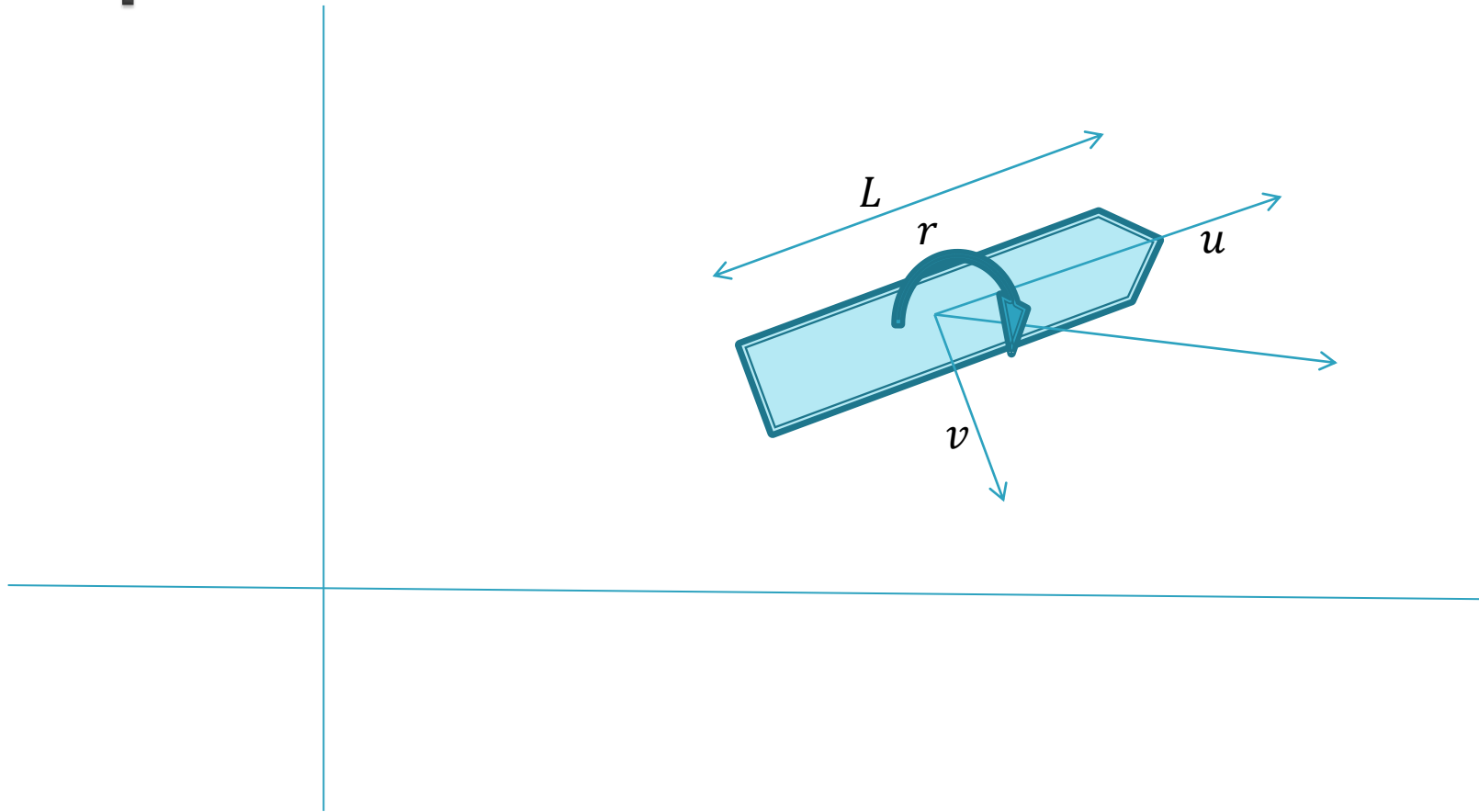
Problem description

- ▶ Describe manoeuvring behaviour of (propelled) ships in terms of
 - Steering parameters
 - rudder angle, thrust
 - design parameters of the ship
 - Mass, length, moment of inertia, water displacement..
- ▶ Focus: stability behaviour for steady manoeuvres (equilibria).

Model description

- ▶ Model of ship with forces...
 - Thrust of the propellor
 - possibly directed by rudder
 - Dissipative friction with the water
 - Coriolis force
 - due to choice of coordinate comoving with the ship
- ▶ Effective masses
 - Effect of dragging along the water by ship
 - Choice of coordinate system
- ▶ Lots and lots of coefficients:
 - functional dependence of forces on velocity

Ship vs earth coordinates



Model equations in rotating ship coordinates

- ▶ $M \dot{\boldsymbol{v}} = \boldsymbol{F} = \boldsymbol{F}^\perp(\boldsymbol{v}) + \boldsymbol{F}_D(\boldsymbol{v}, \alpha) + \boldsymbol{F}_T(\boldsymbol{v}, \alpha)$
- ▶ Where $\boldsymbol{v} = (u, v, Lr)$, and
- ▶ M symmetric mass matrix
- ▶ $\boldsymbol{F}^\perp = (mrv, -mru, 0)$ Coriolis force $\perp \boldsymbol{v}$
- ▶ \boldsymbol{F}_D is a dissipative friction force
- ▶ \boldsymbol{F}_T propulsion force
- ▶ α steering angle

Dissipative force with rudder

$$X_{hull} = \frac{1}{2} \rho L_{pp} T \left(X'_{u'|u'|} |u| |u| + X'_{\beta \gamma} L_{pp} v r \right)$$

$$Y_{hull} = \frac{1}{2} \rho L_{pp} T \left(Y'_{\beta'} |u| v + Y'_{\gamma} L_{pp} u r + Y'_{\beta|\beta|} v |v| + Y'_{\gamma|\gamma|} L_{pp}^2 r |r| \right. \\ \left. + Y'_{\beta|\gamma|} L_{pp} v |r| + Y'_{|\beta|\gamma} L_{pp} |v| r + Y'_{ab} |u^{ay} v^{by}| \text{sign}(v) V^{-ay-by+2} \right)$$

$$N_{hull} = \frac{1}{2} \rho L_{pp}^2 T \left(N'_{\beta} u v + N'_{\gamma} L_{pp} r |u| + N'_{u' \gamma c} L_{pp}^{c_n} |u r^{c_n}| V^{-c_n+1} \text{sign } r \right. \\ \left. + N'_{r|r|} r |r| L_{pp}^2 + N'_{\beta|\beta|} v |v| + N'_{\beta \beta \gamma} r v^2 L_{pp} V^{-1} \right. \\ \left. + N'_{\beta \gamma \gamma} v r^2 L_{pp}^2 V^{-1} \text{sign } u + N'_{ab} |u^{a_n} v^{b_n}| V^{-a_n-b_n+2} \text{sign}(u v) \right)$$

$$X_R = -\frac{1}{2} \rho V_{rr}^{-1} A_R C_L \left(\frac{C_L u_r}{\pi \Lambda} (u_r \sin \delta - v_r \cos \delta)^2 + v_r (u_r \sin \delta - v_r \cos \delta) (u_r \cos \delta + v_r \sin \delta) \right)$$

$$Y_R = \frac{1}{2} (1 + a_H) \rho V_{rr}^{-1} A_R C_L \left(u_r (u_r \sin \delta - v_r \cos \delta) (u_r \cos \delta + v_r \sin \delta) - \frac{C_L v_r}{\pi \Lambda} (u_r \sin \delta - v_r \cos \delta)^2 \right)$$

$$N_R = Y_R x_r - X_R y_r$$

where

$$V_{rr} = \sqrt{u_r^2 + v_r^2}$$

$$u_r = u_p + C_{rue} \left(\sqrt{u_p^2 + \frac{8T_p(u)}{\rho \pi D_p^2}} - u_p \right), \quad \text{case for ahead speed with positive thrust}$$

$$v_r = C_{db} v + C_{dr} x_r r$$

$$u_p = (1 - w) u$$

Thrust with rudder model

- ▶ $X_{prop} = T(u)$
- ▶ $Y_{prop} = 0$
- ▶ $N_{prop} = 0$

- ▶ In the example $T(u)$ decreases as $u \rightarrow \infty$ (?)
 - Power becomes less effective as you move faster through the water.

Model equations in rotating ship coordinates (with fixed thrust)

- ▶ $M \dot{\boldsymbol{v}} = \boldsymbol{F} = \boldsymbol{F}^\perp(\boldsymbol{v}) + \boldsymbol{F}_H(\boldsymbol{v}) + \boldsymbol{F}_T(\alpha)$
- ▶ Where $\boldsymbol{v} = (u, v, Lr)$, and
- ▶ M symmetric mass matrix
- ▶ $\boldsymbol{F}^\perp = (mrv, -mru, 0)$ Coriolis force $\perp \boldsymbol{v}$
- ▶ \boldsymbol{F}_H is a dissipative friction force from Hull
- ▶ \boldsymbol{F}_T propulsion force
- ▶ α steering angle

Thruster model

- ▶ Only keep the Hull force
- ▶ Use fixed thrust in fixed direction to steer.
 - $X = \tau \cos(\alpha)$
 - $Y = \tau \sin(\alpha)$
 - $N = x_r \tau \sin(\alpha)$

 - E.g. take $x_r = -0.5L$

Equilibria

- ▶ Equilibrium is solution with $\dot{\nu} = 0$ i.e.
- ▶ $F = 0$ we are looking for stationary points
 - in the *ships* coordinate system.
 - In *earth* coordinate system: ship moves on straight line or circle with angular velocity $\omega = r + \operatorname{arctg}\left(\frac{v}{u}\right)$ and radius $\frac{\sqrt{u^2+v^2}}{\omega}$

Proof:

- ▶ Earth coordinates (x, y, ϕ) where
 - (x, y) is position of centre of gravity of the ship.
 - ϕ means the orientation angle of the ship

Write $w = u + i v$, and $z = x + i y$, then

$$\dot{z} = w e^{i\phi}$$

$$\dot{\phi} = r + \arg(w) = \omega$$

Now we have $\phi = \omega t + \phi_0$ and

$$z = z_0 + \left(\frac{w e^{i\phi_0}}{i\omega} \right) e^{i\omega t}$$

i.e. the radius of the circular motion is $\frac{|w|}{\omega}$ and angular velocity is ω .

Obvious equilibrium.

- ▶ Both the models have an “obvious” equilibrium at $\alpha = 0$: moving straight ahead
 - $u = \text{constant}(\tau)$
 - $v = 0$
 - $r = 0$
 -

How to study stability

- ▶ Linearise F in a stationary point.
- ▶ Stability determined by Jacobian $DF = \frac{\partial F}{\partial v}$
 - A real 3×3 matrix
 - All depends on eigenvalues:
 - All negative \rightarrow stability
 - 2 real negative, 1 real positive
 - Convergence on a plane, divergence in one direction
 - 1 real negative, 2 negative real part
 - spiralling in, converging to a plane
 - 1 real negative, 2 positive real part
 - Spiralling out converging to plane in one direction
 -

Result

- ▶ Stability depends on design parameters!
 - Not necessarily bad!
- ▶ Example: tune total mass (keep rest fixed)
 - $m' < m'^*$ 3 negative eigenvalue (stability)
 - $m' > m'^*$ 2 negative, 1 positive eigenvalue
- ▶ Explicit model → 1 Obvious negative eigenvalue, what are possibilities for all 3 ?

Fundamental assumption

- ▶ Friction force is dissipative, i.e.

$$\mathbf{v} \cdot F_D(\mathbf{v}, \alpha) < 0$$

- ▶ More precisely: for given thrust and α , if $|\mathbf{v}| \gg 0$ kinetic energy will decrease

$$\frac{dE}{dt} = \frac{1}{2} \frac{d}{dt} (\mathbf{v} \cdot M \mathbf{v}) = M \mathbf{v} = \mathbf{v} \cdot F^\perp + \mathbf{v} \cdot F_D(\mathbf{v}) + \mathbf{v} \cdot F_T < 0$$

Existence of a solution

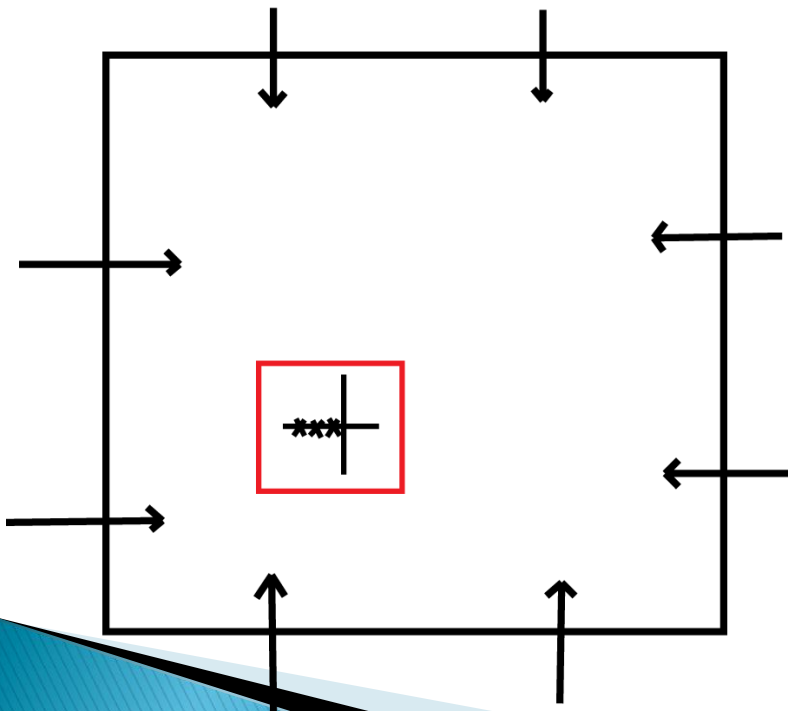
- ▶ From assumptions there *always* exists an equilibrium
 - There exist -1 solution counted with multiplicity.
 - Topological argument

Existence of a solution

- On a large v sphere $v \cdot F < 0$ (inward pointing)
- On the sphere, F is *homotopic* to the vectorfield $-v$ as a map from $S^2 \rightarrow R^3 - 0$ i.e.
 - $(1 - s)F + s(-v) \neq 0$ for all $s \in [0,1]$.
 - This is obvious because $(1 - s)F \cdot v + s(-v) \cdot v < 0$ for all $s \in [0,1]$.
- Therefore $\deg(F) = \deg(-id) = -1$
 - Homotopic maps have the same degree (so it is very robust)
 - To compute the degree of $(-id)$ use the formula on the next slide.
 - The degree really is the map $\deg(F) : \pi_2(S^2) = \mathbf{Z} \rightarrow \pi_2(\mathbf{R}^3 - 0) = \pi_2(S^2) = \mathbf{Z}$ induced on the homotopy groups, and is therefore an integer. It is the two dimensional analog of the winding number. There are many equivalent definitions for smooth maps F , for example, take any volume form μ on S^2 then $\deg(F) = \int_{S^2} F^* \mu / \int_{S^2} \mu$. If you take a volume form concentrated near a generic point then you see the degree is the number of inverse images counted with orientation.

-1 solutions

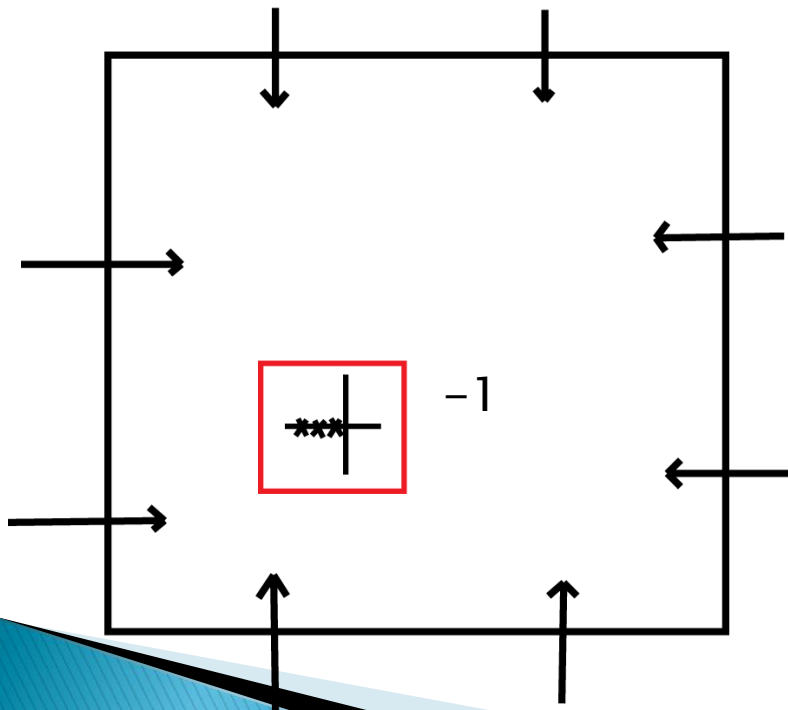
$$\deg(F) = \sum_{F(p)=0} \text{sign}(\det(DF(p)))$$



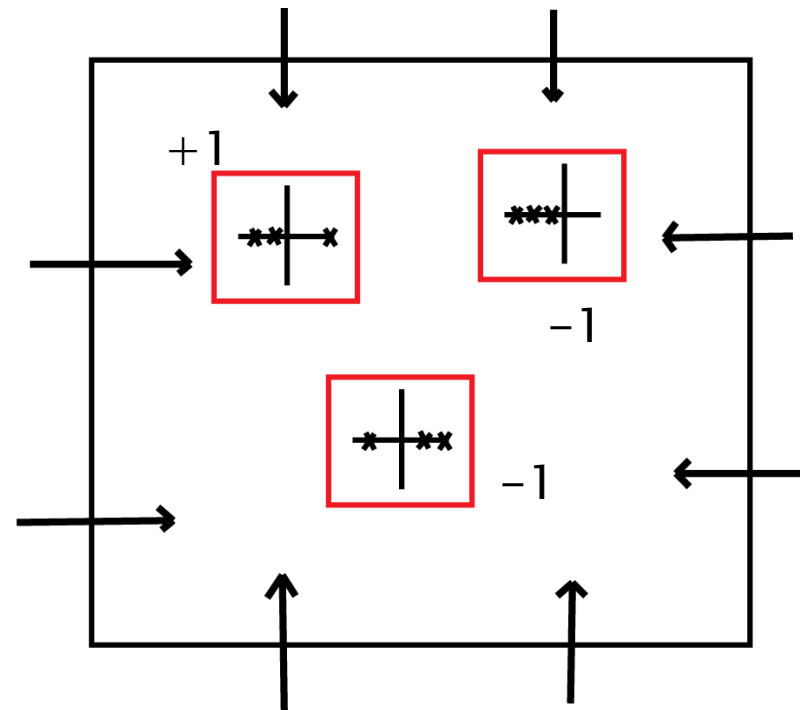
When 1 eigenvalue becomes zero

$$\deg(F) = \sum_{F(p)=0} \text{sign}(\det(DF(p)))$$

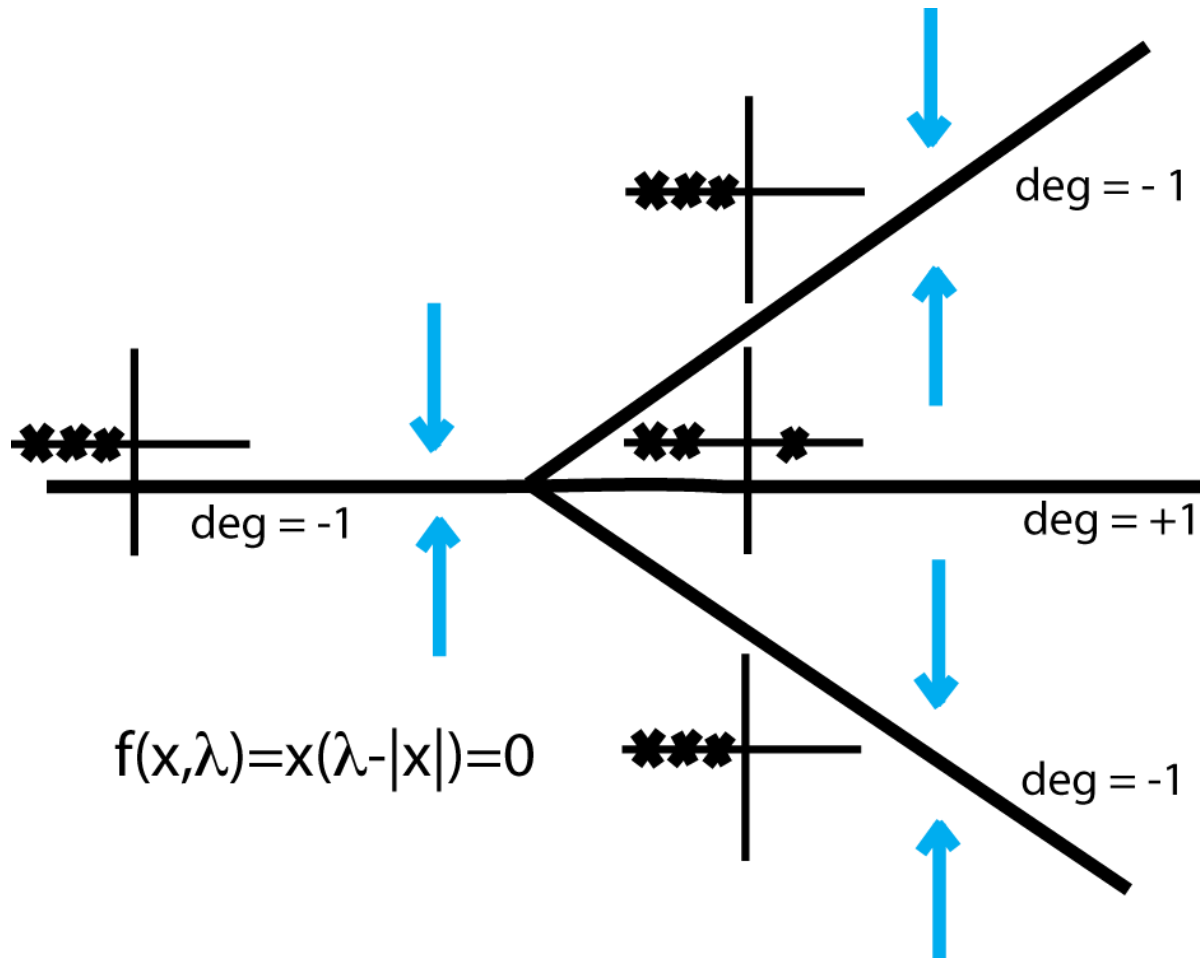
$m < m^*$



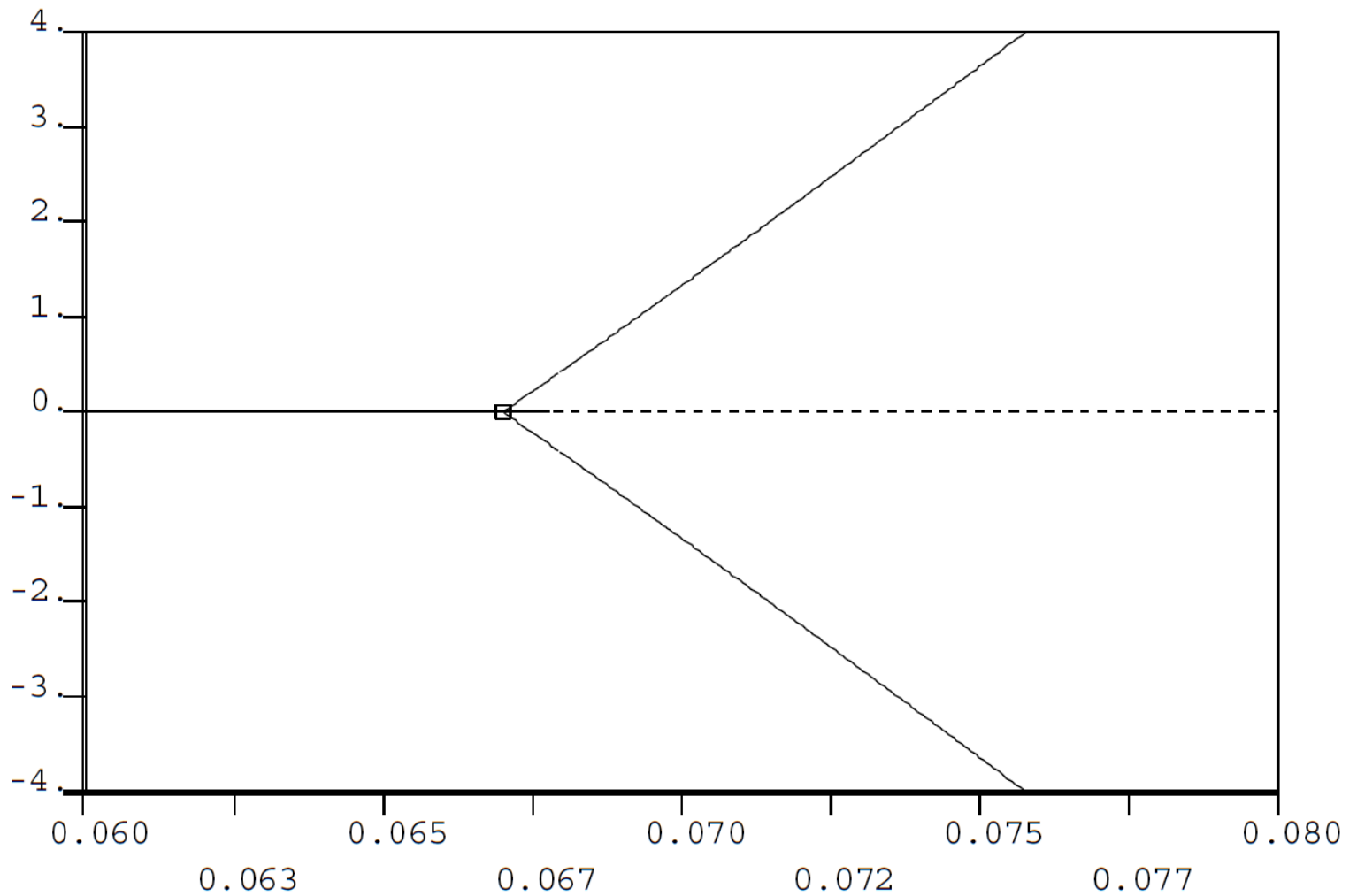
$m > m^*$



Bifurcation diagram

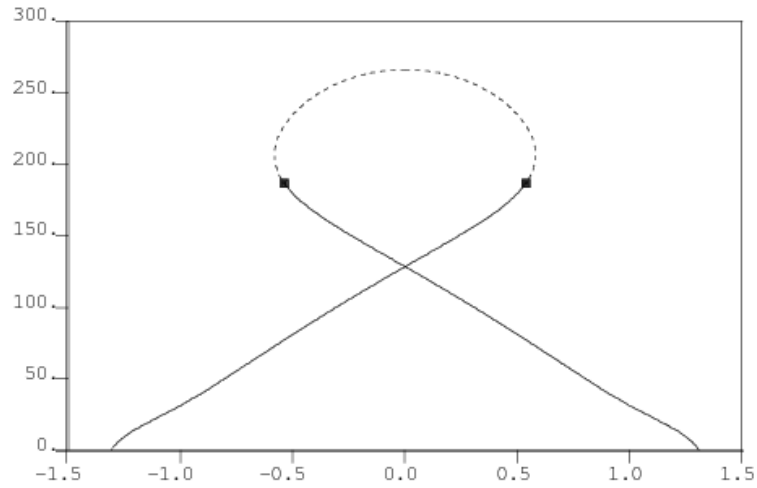


What you find numerically

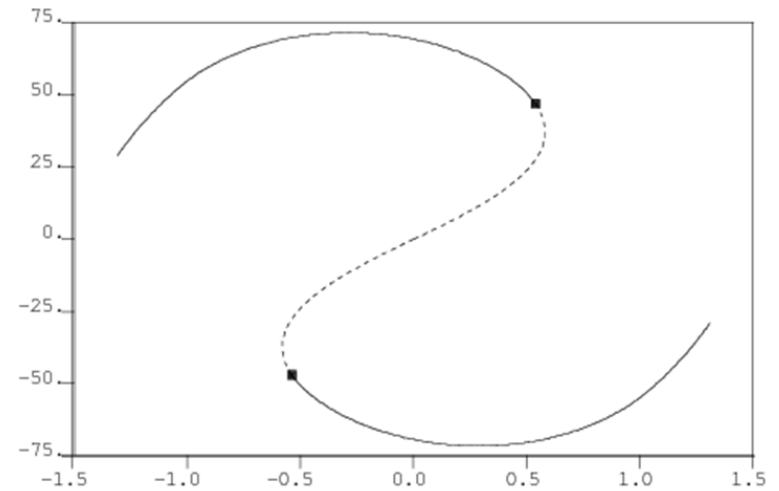


tuning α :

Bifurcation diagrams of equilibria:

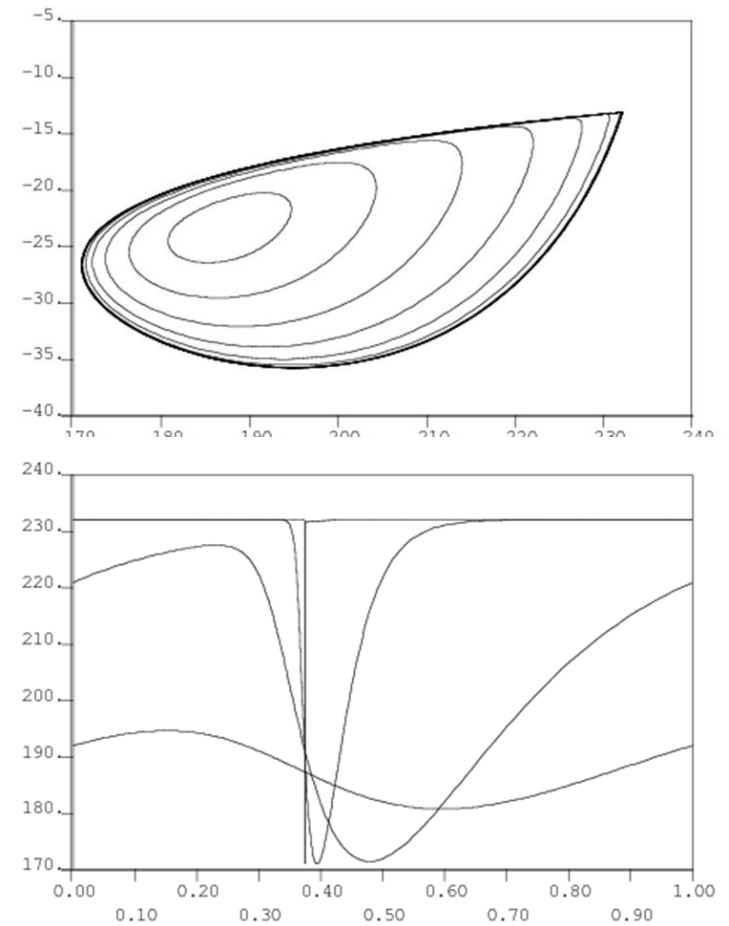
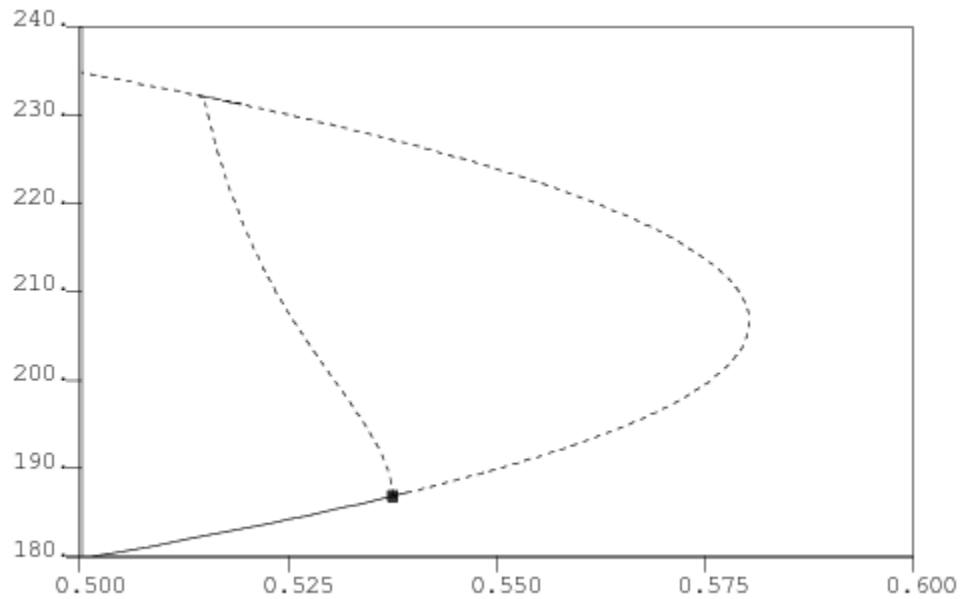


u VS. α



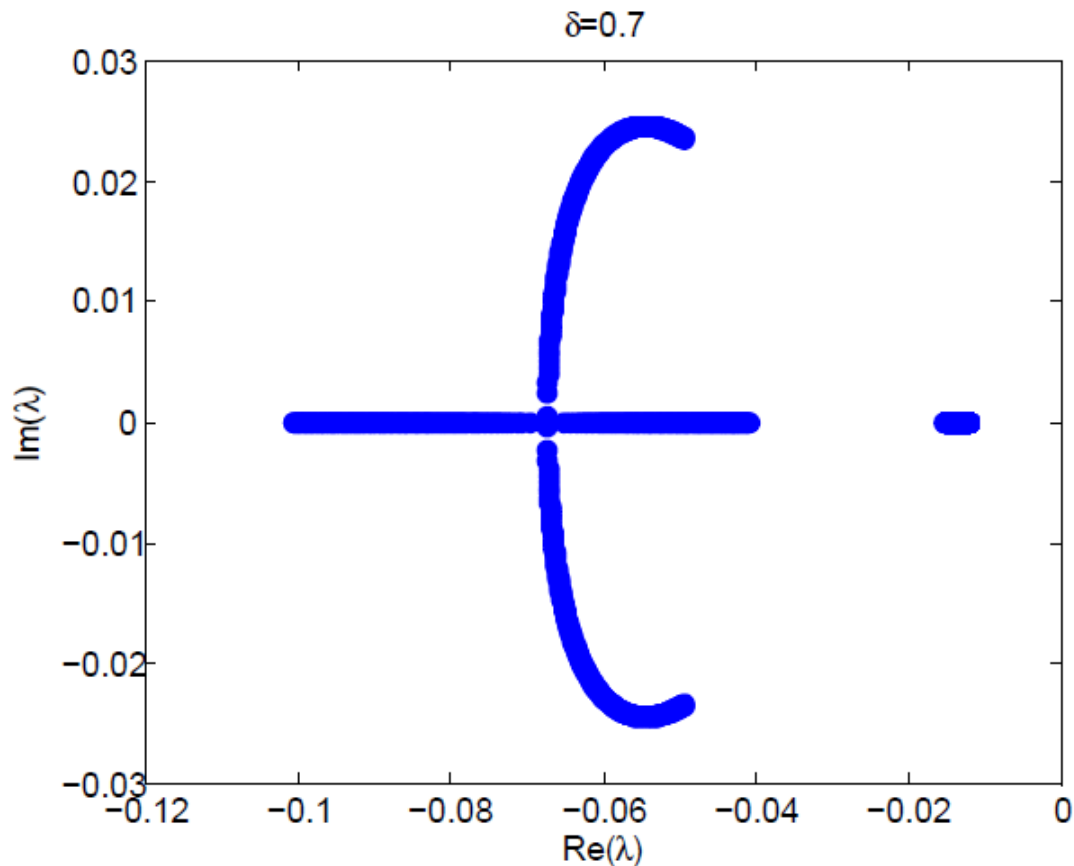
r VS. α

Bifurcation of periodic orbits

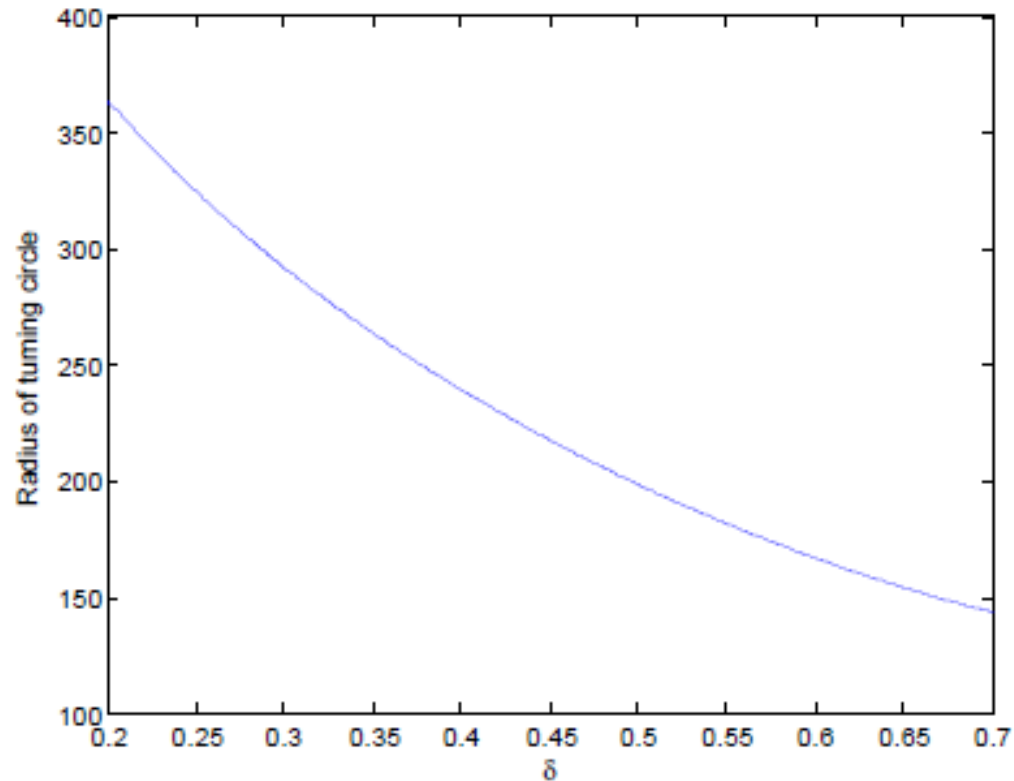


▶ Rudder Model

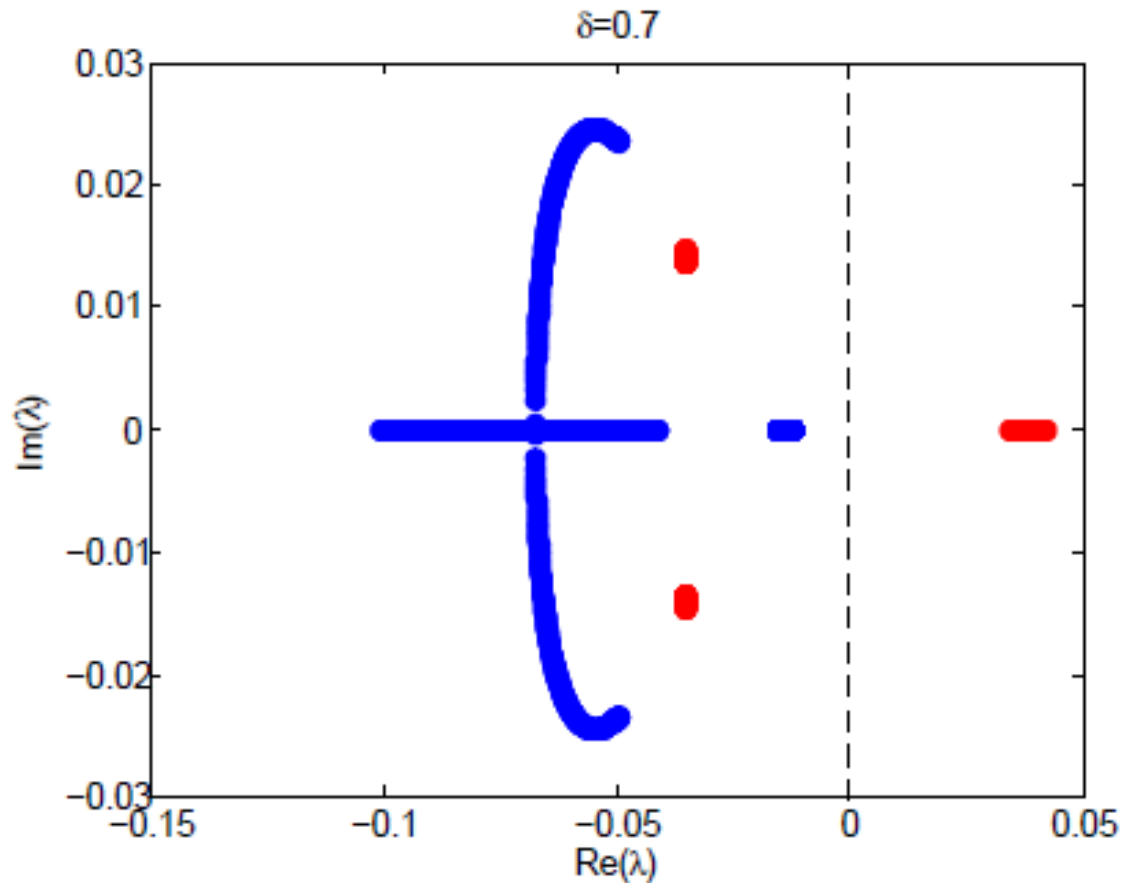
Eigenvalue trajectories for the rudder model



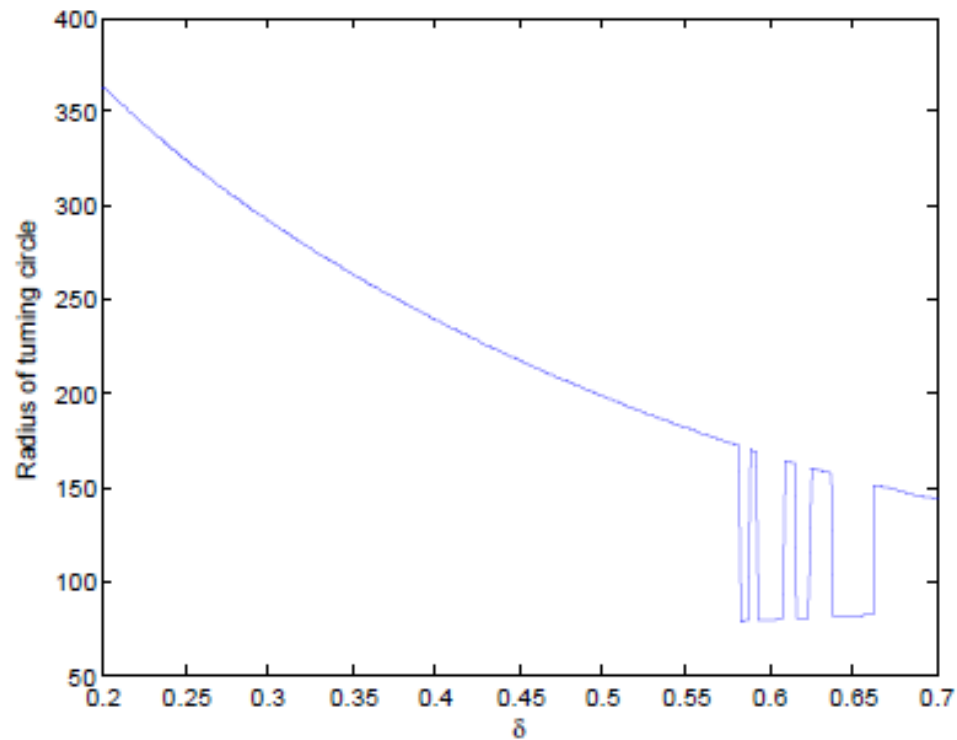
Turning circle vs Rudder angle



Eigenvalue trajectories



Turning circle vs Rudder angle



Outlook

- ▶ Analysis of stationary points and dynamical system point of view, is a useful toolbox.
- ▶ Specialised numerical methods for following stationary points is (fast) alternative to solving full equations.
- ▶ Open question: stabilisation by control.

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